

Laser Communication Requirements Drive Cost-Effective Solutions

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1999 International Symposium on Advanced Radio Technologies
Boulder, Colorado
9 September 1999

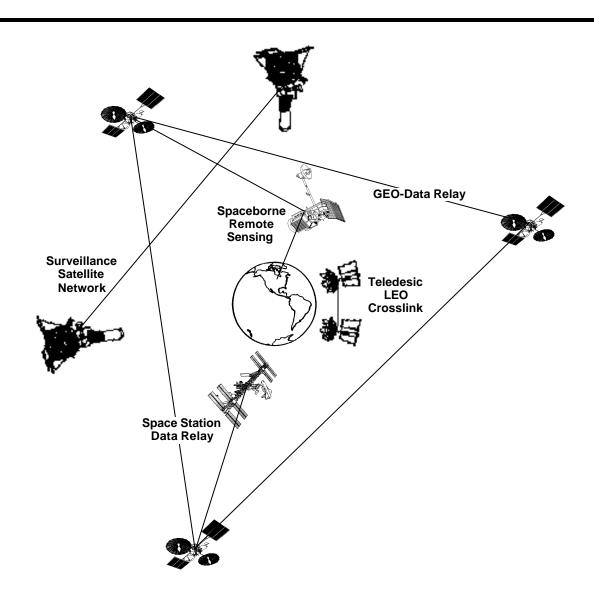


Outline

- Crosslink applications and system requirements
- Benefits of laser communications
- Design trades
- **LEO short-range links**
- **LEO medium-range links**
- LEO and GEO long-range links
- Common and modular hardware solutions
- Development risk
- **○** Summary



Crosslink Applications Support Commercial, Military and Scientific Users





Space-to-Space Applications Stress Different Aspects of the Design

- LEO-to-LEO crosslinks
 - Short range: Up to 1 Gbps over 5000 km
 - Medium range: Up to 7.5 Gbps over 10,000 km
 - Limited attitude knowledge increases acquisition search cone
 - Platform base motion drives residual pointing jitter
- **LEO-to-GEO uplinks**
 - Up to 10 Gbps over 44,000 km
 - Good attitude knowledge reduces GEO search cone
 - LEO platform base motion drives residual pointing jitter
- **○** GEO-to-GEO crosslinks
 - Up to 10 Gbps over 84,000 km
 - Good attitude knowledge reduces acquisition search cone
 - Reduced platform base motion lowers residual pointing jitter

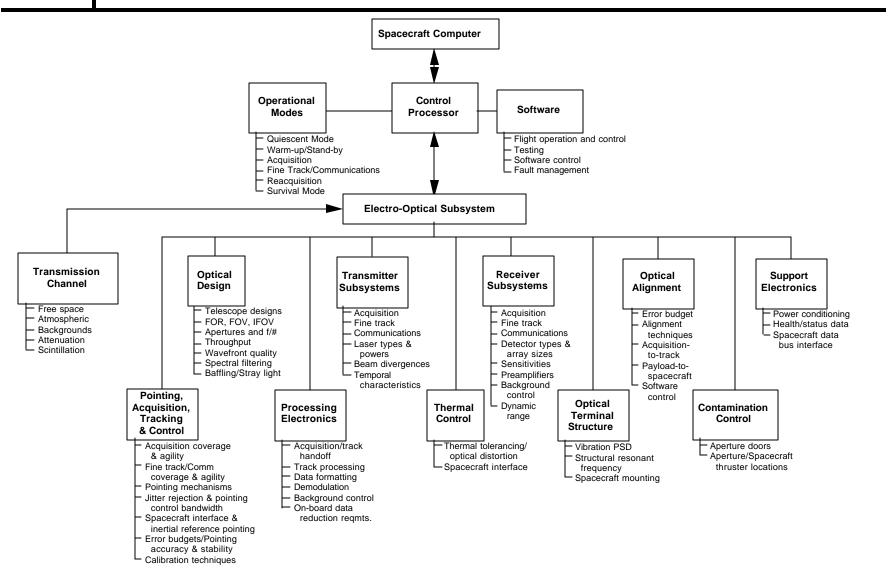


Benefits of Laser Communications

- Increased bandwidth (large carrier frequency)
- Small antenna and swept volume (large antenna gain)
- Secure communication (narrow beam with low probability of intercept and negligible interlink interference)
- **○** Jam resistance (small receive FOV)
- Track and communicate through the sun (small receive FOV)
- Full-duplex communications
- Spectral separation from existing communication systems (non-regulated spectrum)
- Reduced weight and power
- Lower recurring cost



Laser Terminal Comprised of Several Key Subsystems





Key Design Trades Define Cost-Effective Solutions

- Laser transmitter and receiver: 0.85-um AlGaAs Diode with Direct Detection; 1.06-um Nd:YAG with Coherent Detection; 1.55-um EDFA with Direct Detection Preamplifier
- Optical design: On-axis vs. off-axis telescope; Transmit/receive isolation; Solar background rejection; Wavelength multiplexing
- Coarse pointing: Gimbaled telescope vs. gimbaled mirror vs. gimbaled package
- O Acquisition: Scanned beacon and quadrant or array position sensor
- O Fine pointing: High bandwidth steering mirror; High-bandwidth control system; Sensitive position sensors (quadrant vs. nutation)
- O Point-ahead: Accurate orbital data; High-precision point-ahead mirror
- Support electronics: Central processor; Servo control; Thermal control; Power conditioning



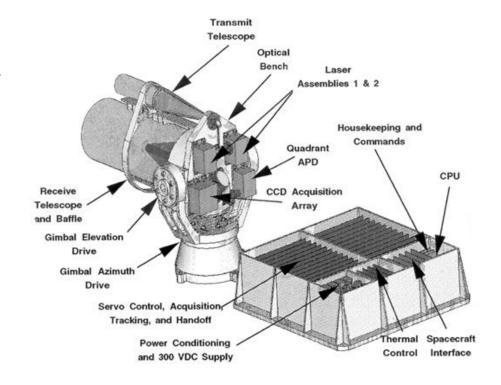
Example Trade: Coarse Pointing Options

Pointing Approach	Pro	Con
Gimbaled Terminal	Smaller laser terminal envelope Fewest optical components that satisfy necessary slew & agility Straightforward telescope baffling Stabilization between laser terminal and platform base motion Near-hemispherical angular coverage	Larger gimbaled mass increases payload mass & power (bearing size, motor size, etc.) Data, electrical, power & some thermal transfer across gimbal Higher pointing-induced momentum & torque
Gimbaled Telescope w/ Coude' Optical Path	Lighter gimbaled mass No thermal or data transfer across gimbal Limited electrical and power transfer across gimbal Straightforward telescope baffling Near-hemispherical angular coverage	Coude' fold mirrors decrease throughput & increase complexity Initial assembly and alignment more complicated Base motion directly imparted to focal plane Moderate pointing-induced momentum & torque
Gimbaled Flat Mirror	Lighter gimbaled mass No thermal or data transfer across gimbal Limited electrical and power transfer across gimbal	Additional large optic Requires larger envelope to separate pointing & fold mirrors More-difficult telescope baffling Stiff structure more difficult to achieve Base motion directly imparted to focal plane Less-than-hemispherical angular coverage Moderate pointing-induced momentum & torque



LEO Short-Range or Low-Data-Rate Applications

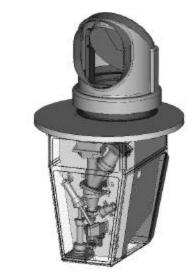
- 10 Mbps at 5000 km range
- 1 @ 75-mW AlGaAs transmitter
- Si CCD acquisition receiver;
 Si quadrant APD track and communications receiver
- O BPPM modulation
- 1.2 inch transmit telescope,4 inch receive telescope
- Gimbaled payload
- 15 lbs, 15 W



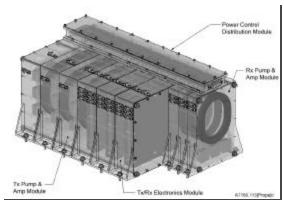


LEO Medium-Range Applications

- → 7.5 Gbps at 6000 km range
- 3 @ 0.3-W EDFA transmitters
- EDFA LNA with 3 @ InGaAs pin photodiode receivers
- OOK modulation
- **7 inch transmit/receive telescope**
- Gimbaled flat mirror
- 25 kg (55 lbs), 100 W



Optical Bench Assembly

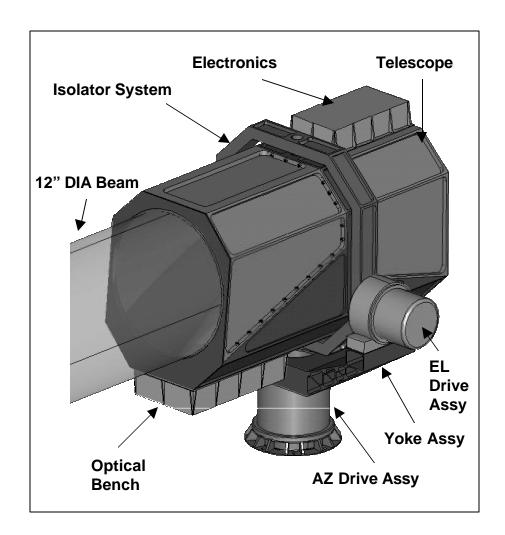


Tx/Rx Electronics Assembly



LEO and GEO Long-Range Applications

- → 10 Gbps at 84,000 km range
- 4 @ 3-W EDFA transmitters
- EDFA LNA with 4 @ InGaAs pin photodiode receivers
- O DPSK modulation
- 12-inch transmit/receive telescope
- Gimbaled telescope
- 90 kg (200 lbs), 300 W



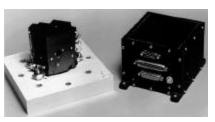


Several Hardware Elements Support a Variety of Applications

- Fine pointing mirror
- O Point-ahead mirror
- Processing electronics
- Optical bench
- Thermal radiator



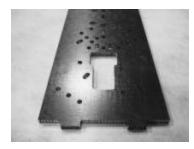
Fine pointing mirror



Point-ahead mirror and electronics



Processing Electronics



Optical Bench



Radiator



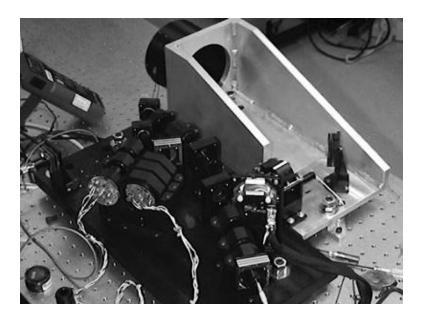
Development Risk Is Manageable

- Laser and receiver reliability and lifetime must be fully verified
- Coarse and fine pointing requirements are met with flight-proven concepts that are modified to meet specific needs
- Electro-optic instrument development is similar to other proven flight systems
 - Optical design for telescope and bench
 - Thermal control
 - Flight processor
 - Software control
 - Heat rejection
 - Radiation effects
 - Contamination control
 - Telemetry and commands
 - Power conditioning
- Cost-effective system integration must be demonstrated



Laboratory Demonstration System Reduces Performance Risk

- Demonstrated 2.5 Gbps simultaneous
 Transmit/Receive operability
- Characterized optical components
- Developed software terminal control and data collection capability
- Demonstrated optical system performance
- Demonstrated performance of fine steering mechanism
- Performed characterization of gimbal
- Demonstrated tracking capability to comply with single-mode fiber coupling requirements
- Demonstrated track loop's ability to reduce pointing errors
- Validated acquisition scenario





Summary

- Laser communications hardware is ready for flight insertion
- Key engineering issues have been conquered on prior electrooptic payloads
 - Optical designs
 - Pointing systems
 - Support electronics
 - Software control
 - Thermal control
- Increased data rate requirements will establish a user pull
 - Communications networks
 - Hyperspectral imaging systems
- Cost control will result after the first system is demonstrated



Acronyms

AlGaAs Aluminum gallium arsenide

APD Avalanche photodiode

O AZ Azimuth

O BPPM Binary pulse position modulation

CCD Charge coupled device

DPSK Differential phase-shift keying

EDFA Erbium-doped fiber amplifier

O EL Elevation

FOR Field of regard

O FOV Field of view

○ FSM Fast steering mirror

GEO Geostationary Earth orbit

O IFOV Instantaneous field of view

O InGaAs Indium gallium arsenide

O LEO Low Earth orbit

O LNA Low-noise amplifier

Nd:YAG
 Neodymium-doped yttrium aluminum garnet

O OOK On-off keying

O PAM Point ahead mirror

○ Tx/Rx Transmit/Receive